

5 Scalar fields

5.1 Kallen-Lehmann representation of a two-point function

Consider for simplicity the case of one particle with mass m and spin 0. Suppose that, at each point x , we can define a hermitian operator $\phi(x)$, a *quantum field*. If we change coordinates, $x \rightarrow x' = \Omega(\Lambda, a)x$, then physical states will change accordingly

$$|\psi\rangle \rightarrow |\psi'\rangle = U(\Lambda, a)|\psi\rangle.$$

What is then $\phi(x')$? We impose it is a scalar field, i.e. its expectation values do not change with coordinates

$$\langle\chi|\phi(x')|\psi'\rangle = \langle\chi|U^{-1}(\Lambda, a)\phi(x')U(\Lambda, a)|\psi\rangle = \langle\chi|\phi(x)|\psi\rangle.$$

This gives the transformation rule

$$\phi(x') = U(\Lambda, a)\phi(x)U^{-1}(\Lambda, a).$$

In particular, note that

$$\phi(x+a) = e^{iP_\mu a^\mu}\phi(x)e^{-iP_\mu a^\mu}.$$

Consider now the vacuum expectation value of the commutator of two scalar hermitian quantum fields at two different space-time points x and y :

$$G(x-y) \equiv \langle 0 | [\phi(x), \phi(y)] | 0 \rangle.$$

The fact that G depends only on the difference $x-y$ is a consequence of the invariance of the vacuum under translations, in fact

$$\langle 0 | \phi(x)\phi(y) | 0 \rangle = \langle 0 | e^{iP \cdot x} \phi(0) e^{-iP \cdot x} e^{iP \cdot y} \phi(0) e^{-iP \cdot y} | 0 \rangle = \langle 0 | \phi(0) e^{-iP \cdot (x-y)} \phi(0) | 0 \rangle.$$

We study then the function

$$G(x) = \langle 0 | [\phi(x), \phi(0)] | 0 \rangle.$$

We insert between the two fields a complete set of states:

$$\mathbb{1} = \sum_n |n\rangle\langle n| = |0\rangle\langle 0| + \int \frac{d^3\vec{p}}{(2\pi)^3 2E_{\vec{p}}} |\vec{p}\rangle\langle \vec{p}| + \dots$$

and we get

$$\begin{aligned} \langle 0 | \phi(x)\phi(0) | 0 \rangle &= \sum_n \langle 0 | \phi(x) | n \rangle \langle n | \phi(0) | 0 \rangle \\ &= \sum_n \langle 0 | e^{iP \cdot x} \phi(0) e^{-iP \cdot x} | n \rangle \langle n | \phi(0) | 0 \rangle = \sum_n e^{-ip_n \cdot x} |\langle 0 | \phi(0) | n \rangle|^2. \end{aligned}$$

This gives

$$\langle 0 | [\phi(x), \phi(0)] | 0 \rangle \sum_n (e^{-ip_n \cdot x} - e^{ip_n \cdot x}) |\langle 0 | \phi(0) | n \rangle|^2 = \int \frac{d^4 q}{(2\pi)^3} (e^{-iq \cdot x} - e^{iq \cdot x}) \tilde{\rho}(q),$$

where

$$\tilde{\rho}(q) = (2\pi)^3 \sum_n \delta^4(q - p_n) |\langle 0 | \phi(0) | n \rangle|^2.$$

We now use the fact that the eigenvalues of P^2 and P^0 are non-negative, and that $\tilde{\rho}(q)$ is invariant under proper orthochronous Lorentz transformations, to recast $\tilde{\rho}(q)$ in terms of a *spectral density* $\rho(q^2)$, as follows:

$$\tilde{\rho}(q) = \rho(q^2) \Theta(q^0).$$

This implies

$$\langle 0 | [\phi(x), \phi(0)] | 0 \rangle = \int \frac{d^4 q}{(2\pi)^3} (e^{-iq \cdot x} - e^{iq \cdot x}) \rho(q^2) \Theta(q^0).$$

We now change variable to $\mu^2 \equiv q^2$, and obtain the *Kallen-Lehmann* representation of the vacuum expectation value of the commutator of two scalar fields at different space-time points

$$\langle 0 | [\phi(x), \phi(0)] | 0 \rangle = \int_0^\infty d\mu^2 \rho(\mu^2) i\Delta(x, \mu^2).$$

The function $\Delta(x, \mu^2)$ is defined as

$$i\Delta(x, \mu^2) = \int \frac{d^4 q}{(2\pi)^3} (e^{-iq \cdot x} - e^{iq \cdot x}) \Theta(q^0) \delta(q^2 - \mu^2) = \int \frac{d^4 q}{(2\pi)^3} e^{-iq \cdot x} \epsilon(q^0),$$

with $\epsilon(q^0) \equiv \Theta(q^0) - \Theta(-q^0)$ the sign of q^0 . The function $\Delta(x, \mu^2)$ has a set of remarkable properties

- (i) $(\square + \mu^2)\Delta(x, \mu^2) = 0$;
- (ii) $\Delta(\Lambda x, \mu^2) = \Delta(x, \mu^2)$ if $\Lambda \in L_+^\uparrow$;
- (iii) $\Delta(x, \mu^2)$ is a function of x^2 and $\epsilon(x^0)$ only;
- (iv) $\Delta(x, \mu^2) = -\Delta(-x, \mu^2)$;
- (v) $\frac{\partial}{\partial x^0} \Delta(x - y, \mu^2)|_{x^0=y^0} = -\delta^3(\vec{x} - \vec{y})$;
- (vi) $\Delta(x, \mu^2) = 0$ if $x^2 < 0$.

Property (vi) ensures that the commutator of two fields at points separated by a space-like distance, i.e. that cannot be connected by a light-ray, vanishes. This ensures the *causality* of the theory.

Let us isolate the contribution of one-particle states

$$|\langle 0|\phi(0)|1\rangle|^2 = \int \frac{d^3\vec{p}}{(2\pi)^3 2E_{\vec{p}}} |\langle 0|\phi(0)|\vec{p}\rangle|^2.$$

Since $\phi(x)$ is a scalar quantum field

$$\langle 0|\phi(0)|\vec{p}\rangle = \sqrt{Z_\phi} \quad (\text{independent of } \vec{p}).$$

This gives

$$(2\pi)^3 \delta^4(q - p_1) |\langle 0|\phi(0)|1\rangle|^2 = \frac{Z_\phi}{2E_{\vec{q}}} \delta(q^0 - E_{\vec{q}}) \Theta(q^0) = Z_\phi \delta(q^2 - m^2) \Theta(q^0).$$

This, and the fact that, for states with more than one particle, $\mu^2 \geq (2m)^2$, gives the final form of the Kallen-Lehmann representation for the vacuum expectation value of the commutator of two scalar fields:

$$\langle 0|[\phi(x), \phi(y)]|0\rangle = iZ_\phi \Delta(x - y, m^2) + \int_{4\mu^2}^{\infty} d\mu^2 \rho(\mu^2) i\Delta(x - y, \mu^2).$$

Suppose that $\phi(x)$ is normalised so as to satisfy *canonical commutation rules*

$$[\dot{\phi}(x), \phi(y)]|_{x^0=y^0} = -i\delta^3(\vec{x} - \vec{y}).$$

We then obtain

$$\begin{aligned} \frac{\partial}{\partial x^0} \langle 0|[\phi(x), \phi(y)]|0\rangle|_{x^0=y^0} &= -i\delta^3(\vec{x} - \vec{y}) \\ &= \left(Z_\phi + \int_{4\mu^2}^{\infty} d\mu^2 \rho(\mu^2) \right) - i\delta^3(\vec{x} - \vec{y}). \end{aligned}$$

This gives the normalisation condition

$$1 = Z_\phi + \int_{4\mu^2}^{\infty} d\mu^2 \rho(\mu^2).$$

Since $\rho(\mu^2)$ is non-negative, and Z_ϕ is by construction non-negative, we obtain the constraint $0 \leq Z_\phi \leq 1$. If $Z_\phi = 1$, then the field $\phi(x)$ creates states that do not overlap with two- or more-particle states. It is therefore a *free field*.

Infinitesimal transformations.

5.2 Free scalar fields

Consider now a free quantum scalar field $\phi(x)$, satisfying

$$U(\Lambda, a)\phi(x)U^{-1}(\Lambda, a) = \phi(\Lambda x + a),$$

with the additional condition

$$\langle 0|\phi(0)|\vec{p}\rangle = 1 \implies \langle 0|\phi(x)|\vec{p}\rangle = e^{-ip \cdot x}.$$

The effect of this field is to transform $|\vec{p}\rangle$ into the vacuum. We then construct an *annihilation operator* $a(\vec{p})$ that “destroys” a particle of momentum \vec{p} , as follows:

$$\begin{aligned} a(\vec{p})|0\rangle &= 0, \\ a(\vec{p})|\vec{p}'\rangle &= (2\pi)^3 2E_{\vec{p}} \delta^3(\vec{p} - \vec{p}')|0\rangle. \end{aligned}$$

Given this operator, a free hermitian scalar quantum field $\phi(x)$ is given by

$$\phi(x) = \int \frac{d^3\vec{p}}{(2\pi)^2 2E_{\vec{p}}} (e^{-ip \cdot x} a(\vec{p}) + e^{ip \cdot x} a^\dagger(\vec{p})).$$

From the definition of an adjoint operator, we get that $a^\dagger(\vec{p})$ “creates” a particle of momentum \vec{p} , as follows

$$a^\dagger(\vec{p})|0\rangle = |\vec{p}\rangle,$$

and is therefore called *creation operator*. We can construct two-particle states by successive application of creation operators

$$a^\dagger(\vec{p}_1)a^\dagger(\vec{p}_2)|0\rangle = a^\dagger(\vec{p}_1)|\vec{p}_2\rangle = |\vec{p}_1, \vec{p}_2\rangle.$$

How is $|\vec{p}_1, \vec{p}_2\rangle$ related to $|\vec{p}_2, \vec{p}_1\rangle = a^\dagger(\vec{p}_1)a^\dagger(\vec{p}_2)|0\rangle$? This depends on the relation between $a^\dagger(\vec{p}_1)a^\dagger(\vec{p}_2)$ and $a^\dagger(\vec{p}_2)a^\dagger(\vec{p}_1)$. If we impose commutation relations

$$[a(\vec{p}), a(\vec{p}')] = 0, \quad \text{and} \quad [a(\vec{p}), a^\dagger(\vec{p}')] = (2\pi)^3 2E_{\vec{p}} \delta^3(\vec{p} - \vec{p}'),$$

we obtain $\langle 0|[\phi(x), \phi(y)]|0\rangle = 0$ if $(x - y)^2 < 0$, which preserves causality. These commutation relations imply

$$|\vec{p}_2, \vec{p}_1\rangle = |\vec{p}_1, \vec{p}_2\rangle,$$

which implies these particles are *bosons*.

We now construct an operator that gives us the total energy of a state $|\vec{p}_1, \vec{p}_2, \dots, p_N\rangle$. This is the Hamiltonian

$$\begin{aligned} H &= \int \frac{d^3\vec{p}}{(2\pi)^2 2E_{\vec{p}}} E_{\vec{p}} a^\dagger(\vec{p})a(\vec{p}), \\ H|\vec{p}_1, \vec{p}_2, \dots, p_N\rangle &= \sum_{i=1}^N E_{\vec{p}_i} |\vec{p}_1, \vec{p}_2, \dots, p_N\rangle \implies \langle \psi|H|\psi\rangle \geq 0. \end{aligned}$$

Also, a direct calculation yields

$$[H, \phi(x)] = -i\partial_0\phi(x) \implies H = P_0 \text{ (generator of time translations).}$$

Similarly, we can construct the *momentum operator*:

$$\vec{P} = \int \frac{d^3\vec{p}}{(2\pi)^2 2E_{\vec{p}}} \vec{p} a^\dagger(\vec{p}) a(\vec{p}).$$

If we build the operator $P^\mu \equiv (H, \vec{P})$, we obtain $[P^\mu, \phi(x)] = -i\partial^\mu\phi(x)$, which means that P^μ is the generator of translations. Similarly, $J_{\mu\nu} = x_\mu P_\nu - x_\nu P_\mu$ is the generator of Lorentz transformations, and since these particles have spin zero, there is no other contribution to $J_{\mu\nu}$. Consistently, $W_\mu = 0$. Note that this representation of fields is appropriate for both massive and massless particles.

A direct computation shows that the Hamiltonian H can be expressed in terms of a Hamiltonian density $\mathcal{H}(x)$ as follows:

$$H = \int d^3\mathcal{H}(x), \quad \mathcal{H}(x) = \frac{1}{2}\dot{\phi}^2 + \frac{1}{2}(\vec{\nabla}\phi)^2 + \frac{1}{2}m^2\phi^2 \dots$$

where the dots give a constant contribution to the energy, which we can neglect. Since $\mathcal{H}(x)$ can be expressed in terms of local fields

$$[\phi(x), \phi(y)] = 0, \quad (x - y)^2 < 0 \implies [\mathcal{H}(x), \mathcal{H}(y)] = 0, \quad (x - y)^2 < 0.$$

Therefore, Hamiltonian densities made of local operators are causal.

Last, we observe that $\phi(x)$ satisfies the equation $(\square + m^2)\phi(x) = 0$. But this is the classical equation of motion obtained from the most general Lorentz-invariant quadratic Lagrangian for a classical scalar field:

$$\mathcal{L} = c_1(\partial_\mu\phi)(\partial^\mu\phi) + c_2\phi^2,$$

by computing the corresponding equations of motions, and setting

$$c_1 = \frac{1}{2}, \quad c_2 = -\frac{1}{2}m^2.$$

In fact, one could start from \mathcal{L} , solve the corresponding equation of motion, and promote the solution $\phi(x)$ to an operator by introducing the operators of creation and annihilation, and interpreting their action as before. Such a procedure is called *canonical quantisation*. Note that, if we wish to interpret the solutions of a classical Lagrangian as a quantum field, we obtain important constraints on the coefficients c_1 and c_2 . In fact, from the Lagrangian, we can compute the Hamiltonian density as follows

$$\pi(x) \equiv \frac{\partial\mathcal{L}}{\partial(\partial_0\phi(x))}, \quad \text{and} \quad \mathcal{H} = \pi\dot{\phi} - \mathcal{L}.$$

In the present case

$$\pi(x) = \dot{\phi}, \quad \text{and} \quad \mathcal{H} = c_1\pi^2 + c_1(\vec{\nabla}\phi)^2 - c_2\phi^2.$$

If the term containing π^2 is to be interpreted as the kinetic energy, we must have $c_1 > 0$. Also, if we want that the energy is bounded from below, we need to have $c_2 < 0$. The actual value $c_2 = -m^2 c_1$ is set by imposing that this quantised Hamiltonian is the energy of particles of mass m .